



Tomographic Study of Bars from N-body Simulations

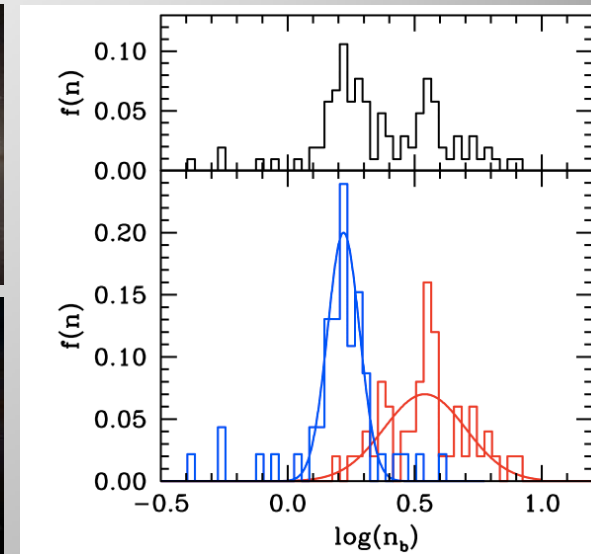
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Bulges in Disk Galaxies

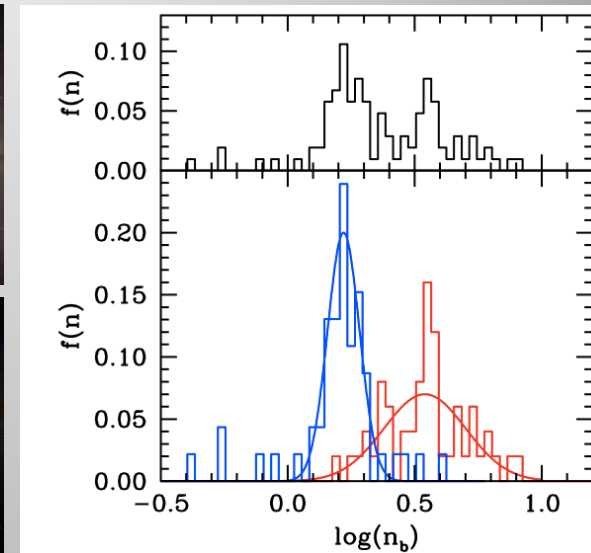
- Classical Bulge
 - Mini-ellipticals
 - Large Sérsic index ($n > 2$)
 - Merger or dissipational process
- Pseudo-Bulge
 - Disk-like
 - Small Sérsic index ($n < 2$)
 - Secular evolution



Fisher & Drory (2008)

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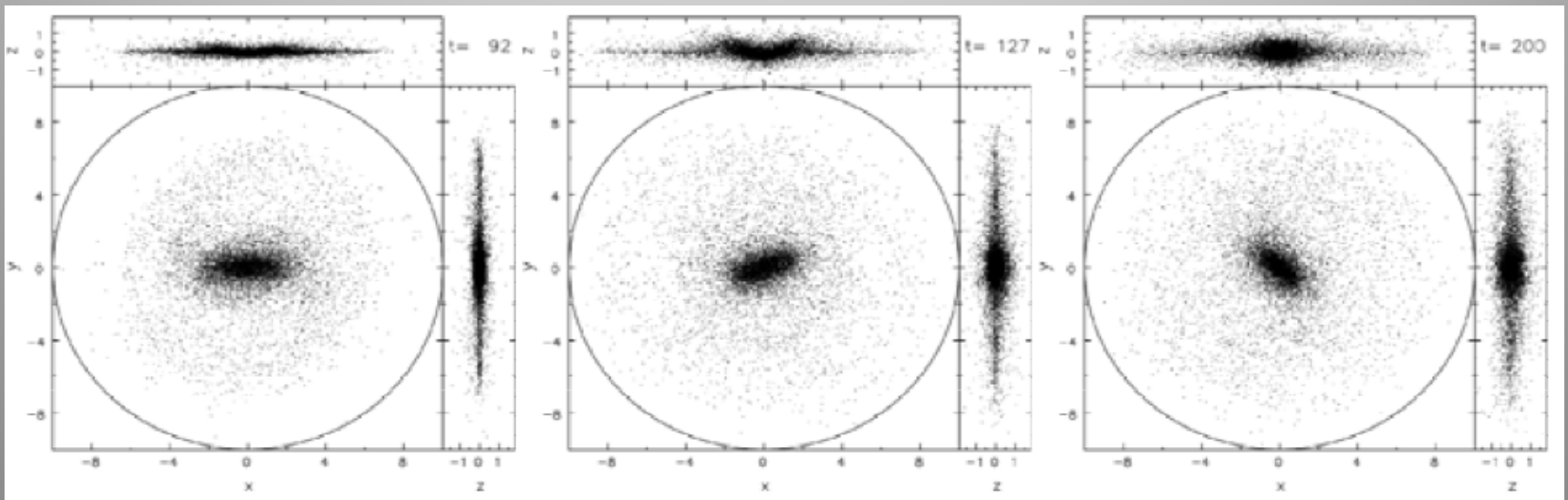
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 - Connection with bars (Burbidge & Burbidge 1959; Jarvis 1986; Shaw 1987; Bureau & Freeman 1999; Lutticke et al. 2000; Bureau & Athanassoula 2005)



Buckling “Fire-hose” Instability of the Bar

- A dynamical instability of thin or elongated galaxies found in 3D N-body simulations (Combes & Sanders 1981)
- Cause the inner region of the bar to puff up in the vertical direction (Combes et al. 1990; Raha et al. 1991; Merritt & Sellwood 1994; Athanassoula & Misiriotis 2002; Patsis et al. 2002; O’Neill & Dubinski 2003; Martinez-Valpuesta & Shlosman 2004)



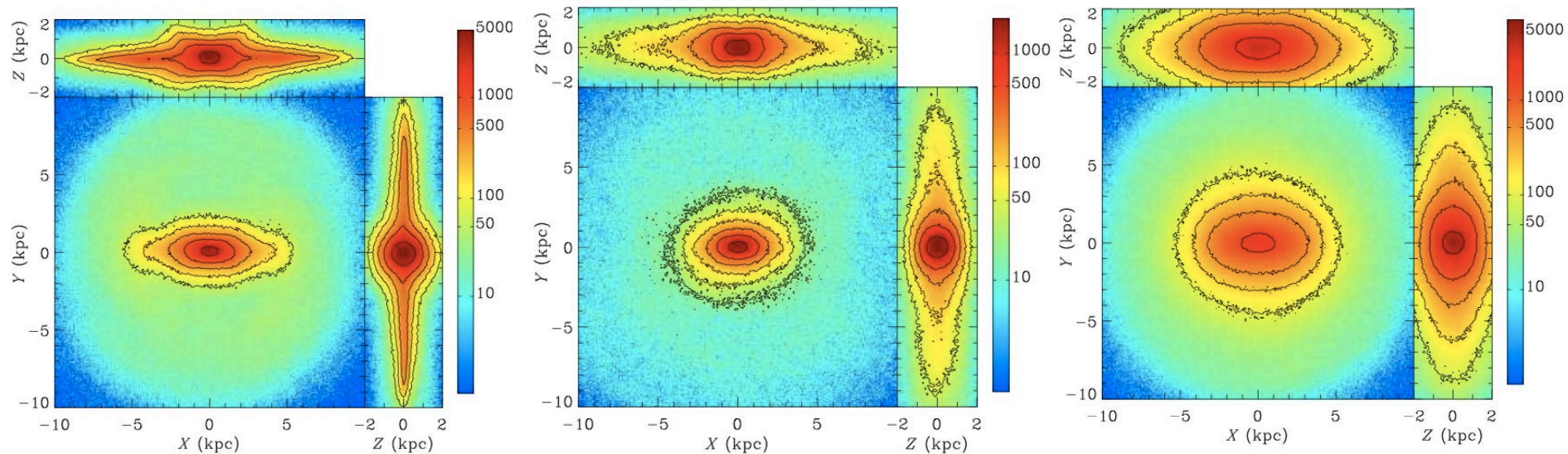
Bar formation → Buckling instability → Saturation → B/PS bulges

Questions Not Well Understood

- What is the density distribution and kinematic properties of the boxy/peanut-shaped bulge in the face-on view?
- Do the properties of the boxy/peanut-shaped bulge depend on the buckling strength of the bar?
- Does the bar have two components, i.e., the boxy/peanut-shaped bulge and extended thin component?
- What are the kinematic properties of particles inside the X shape related to the peanut structure?

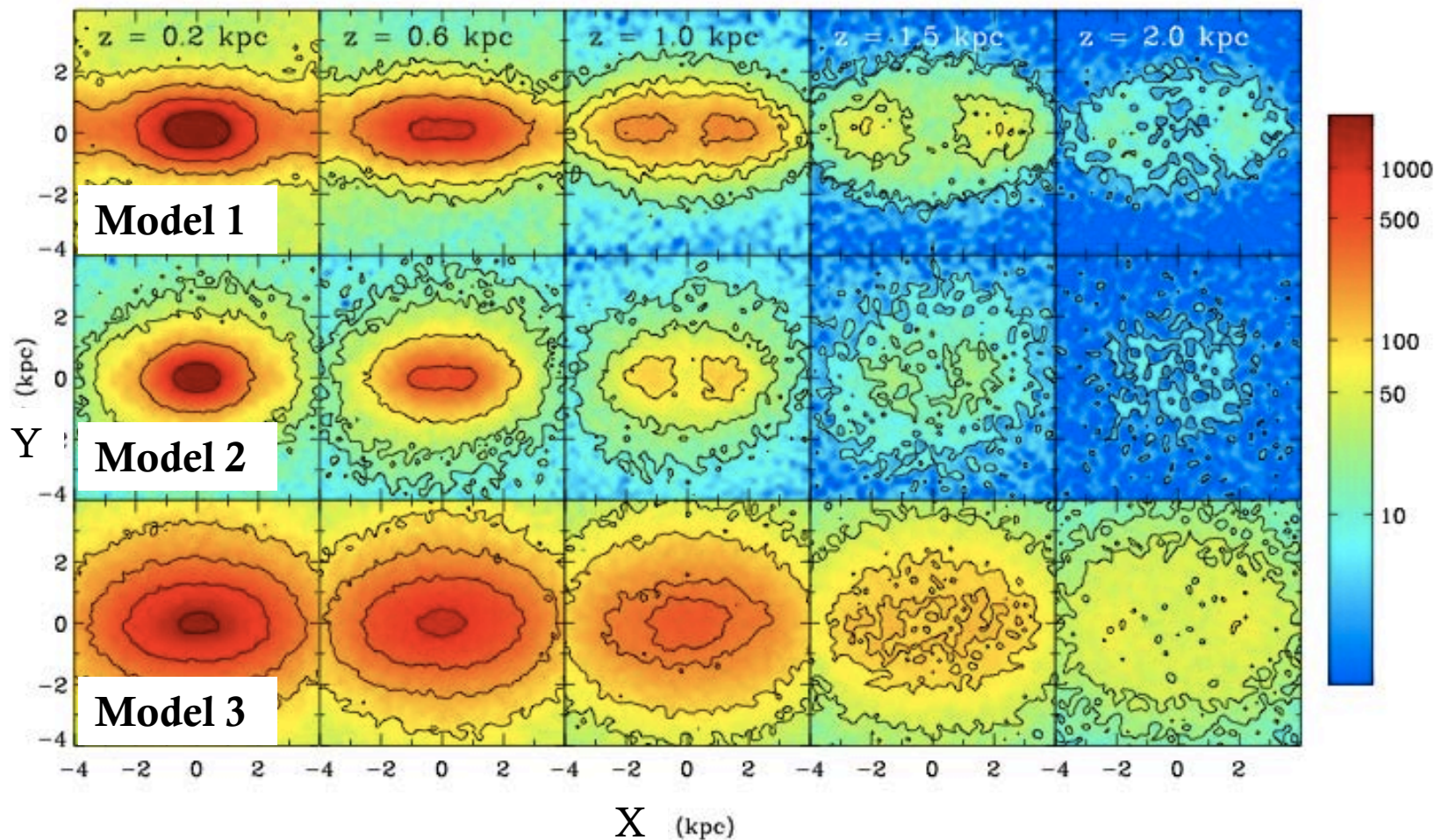
N-body Simulation

- Three models with different buckling strength
- Model 1: thin disk, live halo, strongly buckled
- Model 2: thin disk, rigid halo, buckled
- Model 3: thick disk, rigid halo, weakly buckled



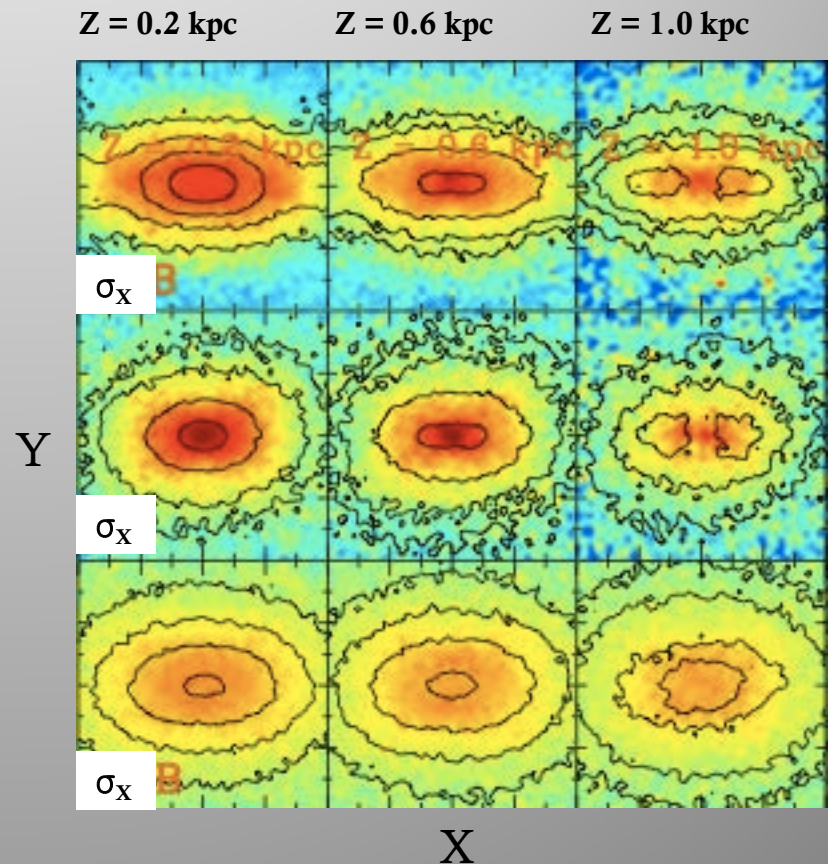
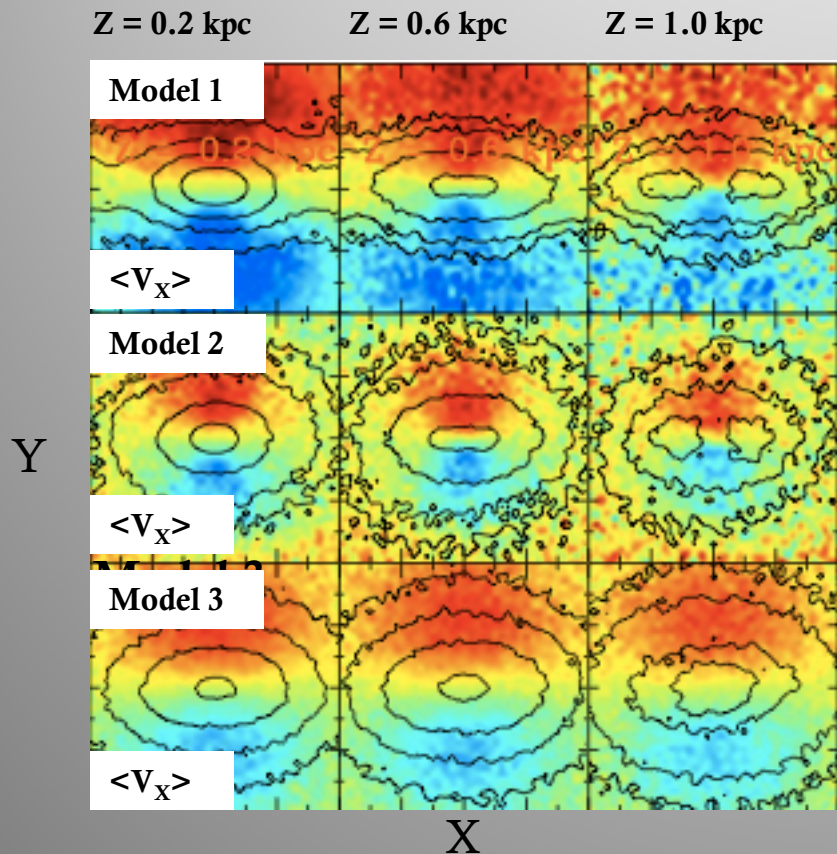
Slices perpendicular to the Z-axis

- Density maps at different heights



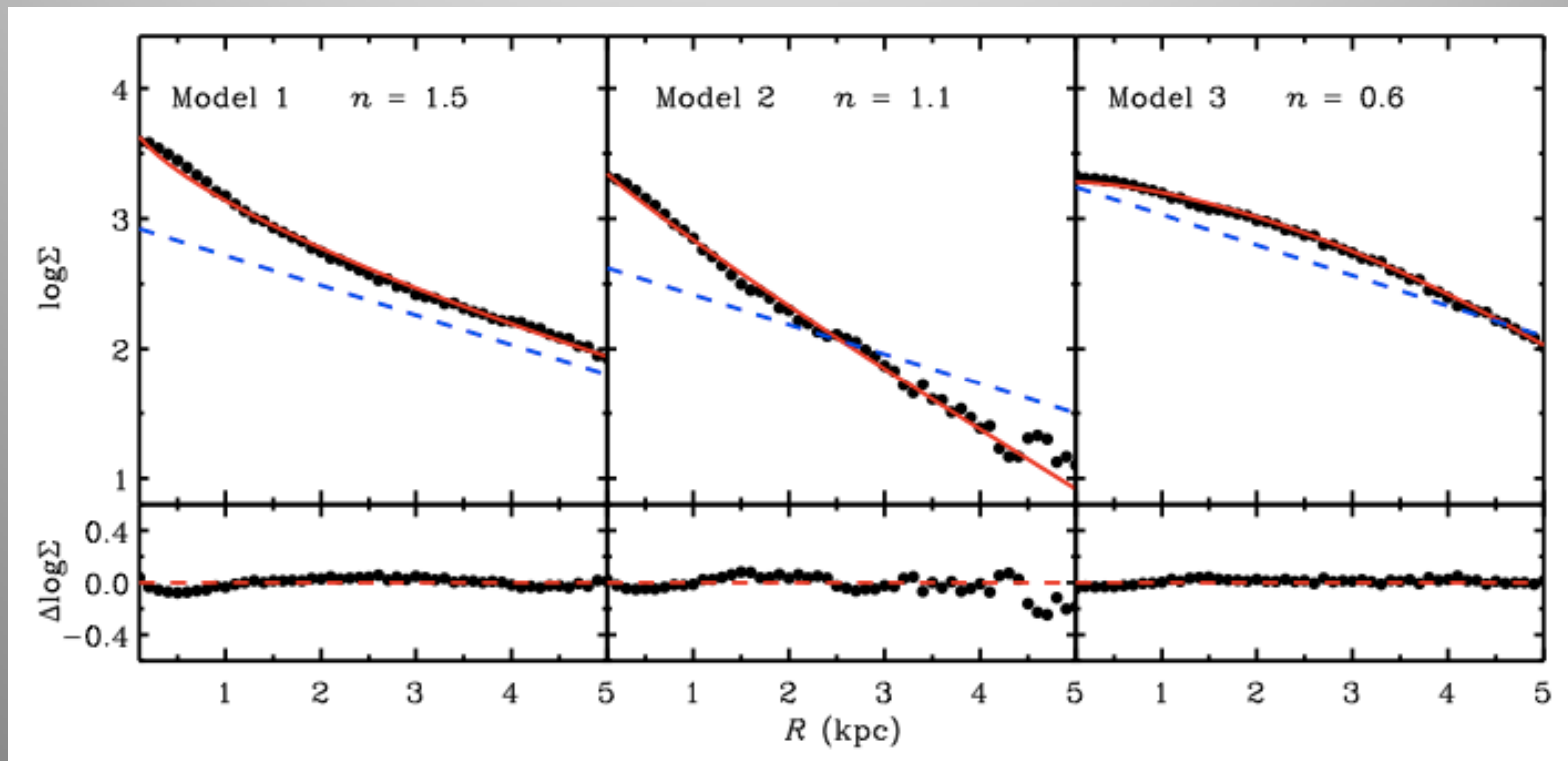
Slices perpendicular to the Z-axis

- Average velocity maps and velocity dispersion maps
- Different kinematic properties within different regions



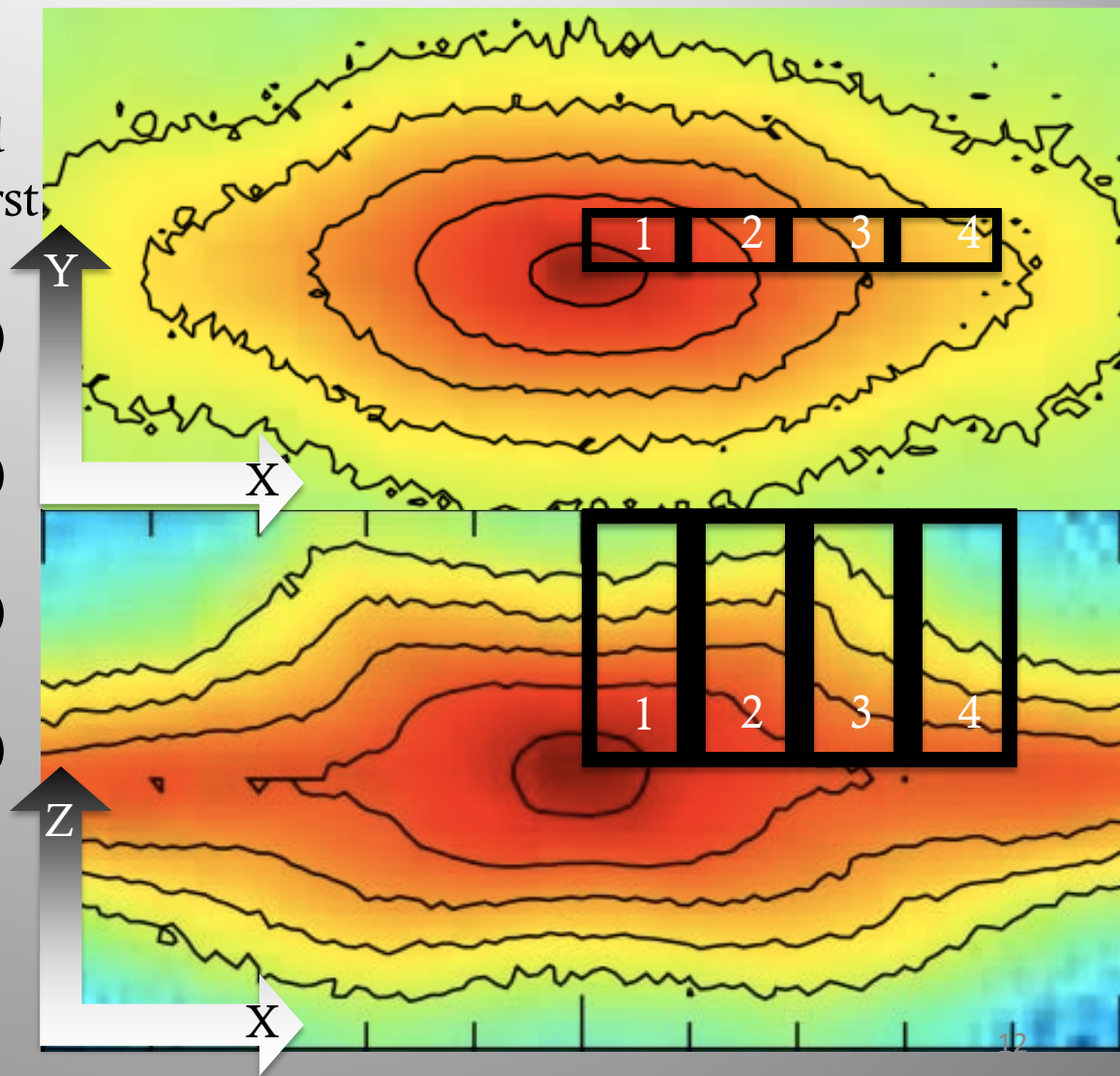
Density Profiles of the B/PS Bulge

- Face-on density profile along the major axis of the bar
 - Well described with a single Sérsic function, with larger index for strongly buckled bar (~ 1.5) than for the weakly buckled bar (~ 0.6)
 - No evidence for two components within the bar region



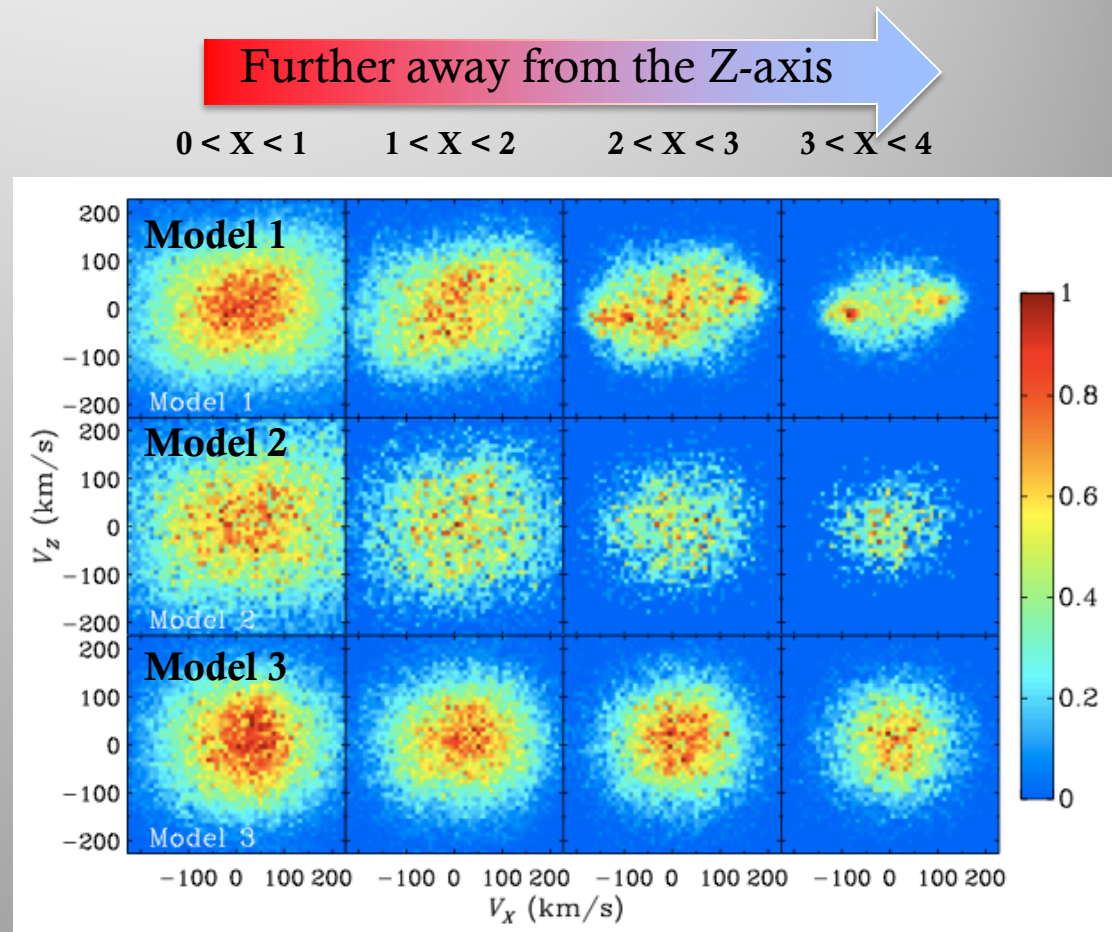
V_Z -- V_X Diagram

- 3D boxes with $Z > 0$ kpc and $0 \text{ kpc} < Y < 0.5 \text{ kpc}$ in the first octant space
- $0 \text{ kpc} < X < 1 \text{ kpc}$ (Region 1)
 - Close to center
- $1 \text{ kpc} < X < 2 \text{ kpc}$ (Region 2)
 - Inner edge of the peanut
- $2 \text{ kpc} < X < 3 \text{ kpc}$ (Region 3)
 - Outer edge of the peanut
- $3 \text{ kpc} < X < 4 \text{ kpc}$ (Region 4)
 - Thin bar region



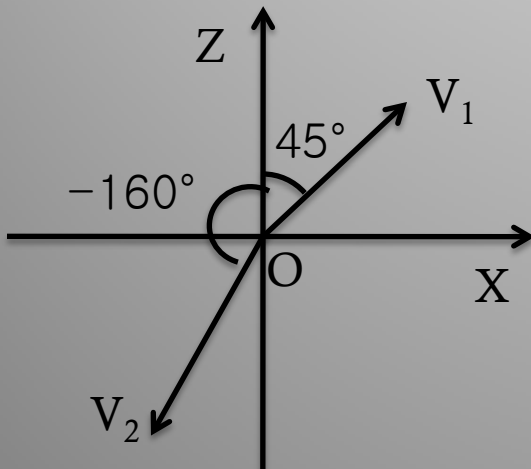
V_Z -- V_X Diagram

- Region 1
 - Large vertical motion, weak positive slope
- Region 2
 - Large vertical motion, positive slope
- Region 3
 - Small vertical motion, weak positive slope
- Region 4
 - Very small vertical motion, flat slope



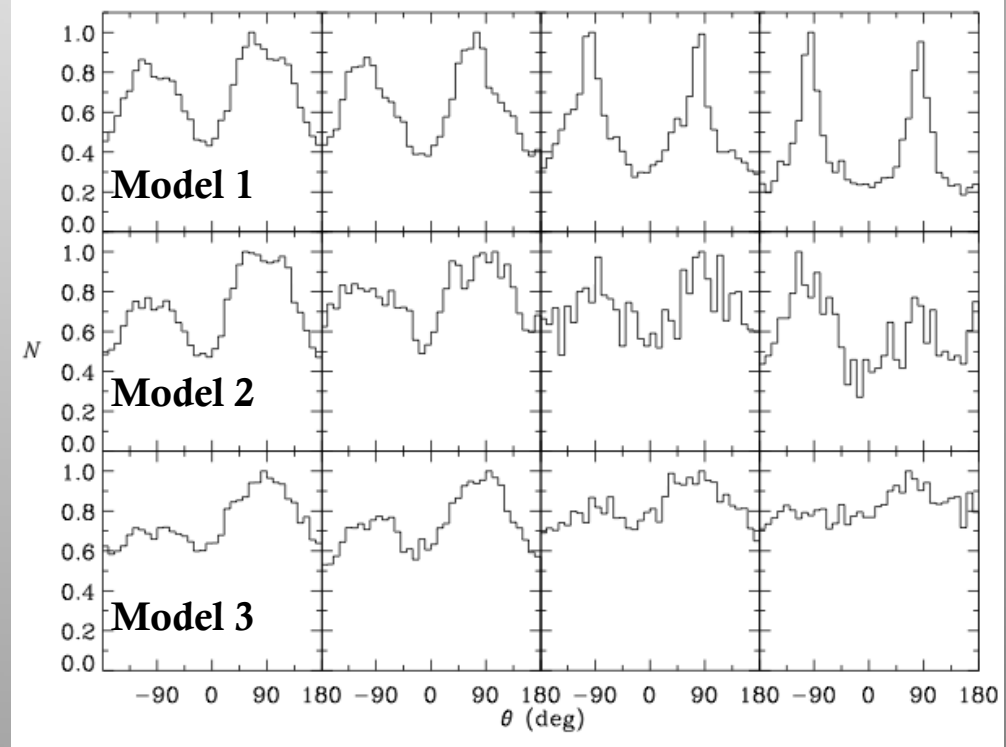
Distributions of Particle Moving Direction

- Different regions with different peak positions, indicating the prevalence of particles on different orbits
 - Inner regions: peak less than 90° or -90°
 - Outer regions: peak at $\pm 90^\circ$ with small dispersion or weak amplitude



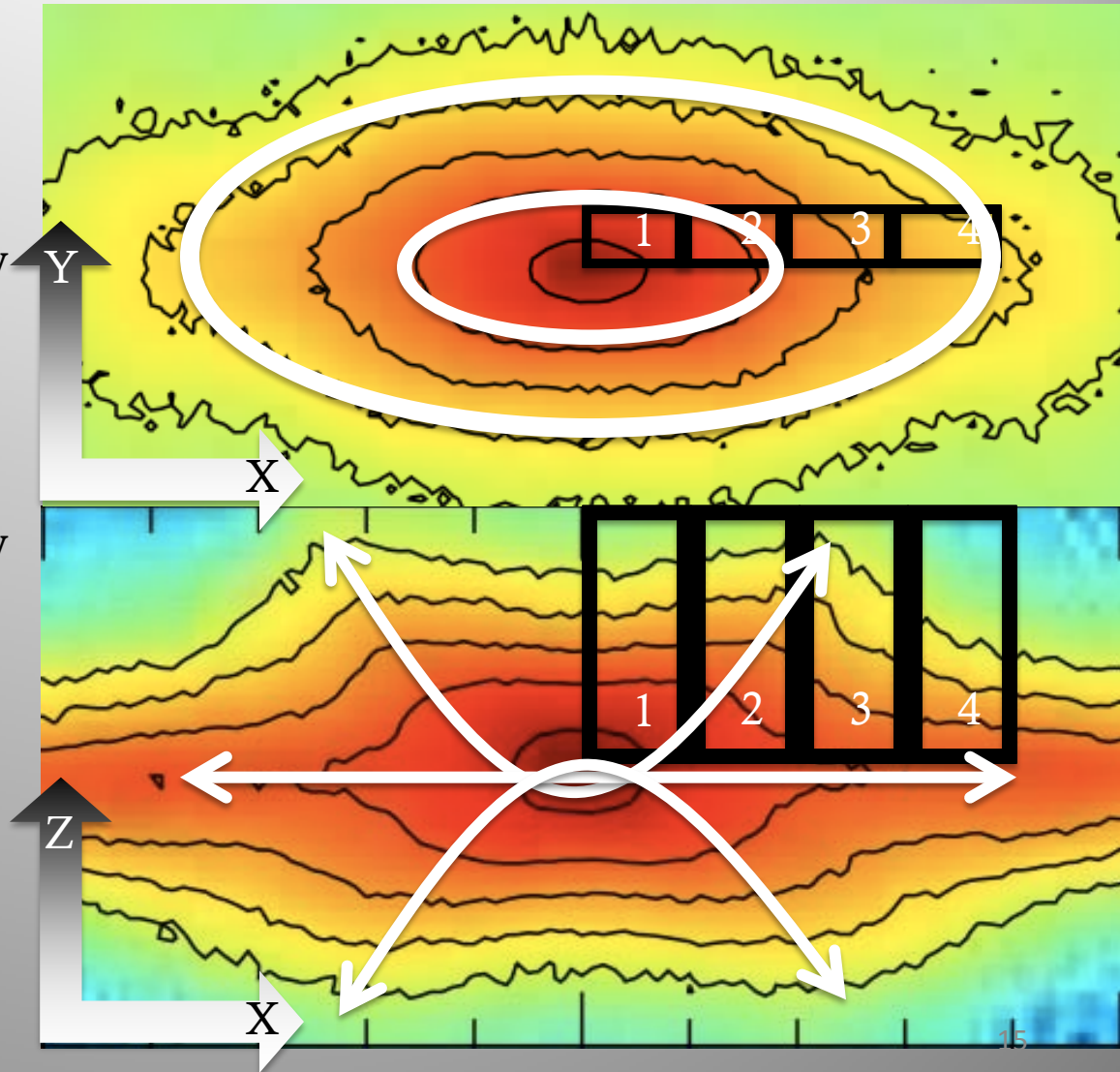
Further away from the Z-axis

$0 < X < 1$ $1 < X < 2$ $2 < X < 3$ $3 < X < 4$

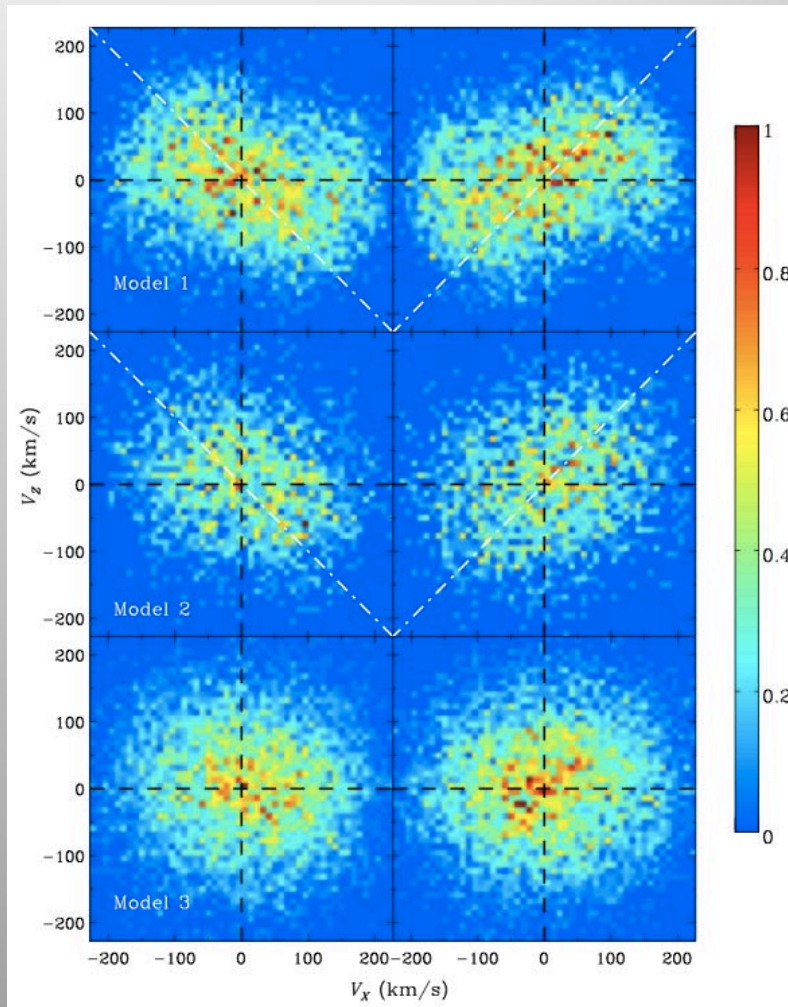
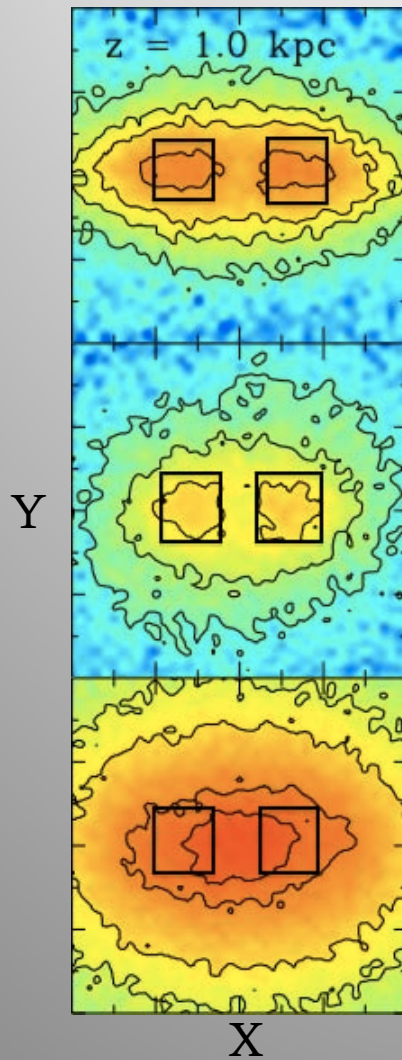


Implications for the Orbits

- Inner regions
 - Bifurcated x1 orbits
 - Ellipse-like in face-on view
 - Banana shaped in edge-on
- Outer regions
 - Regular x1 orbits
 - Ellipse-like in face-on view
 - Little vertical perturbation
- The relative importance of the two orbits produces the observed shape.

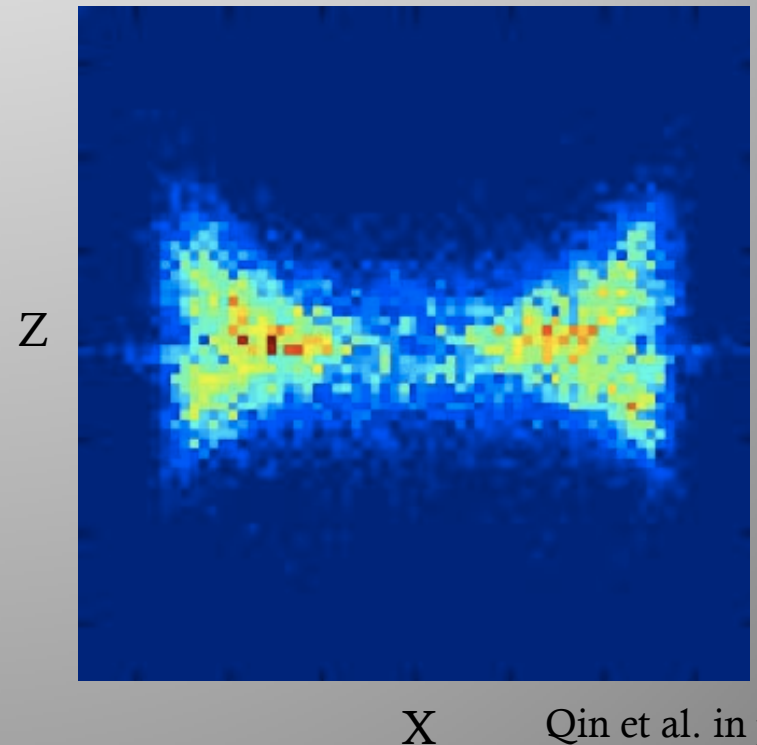
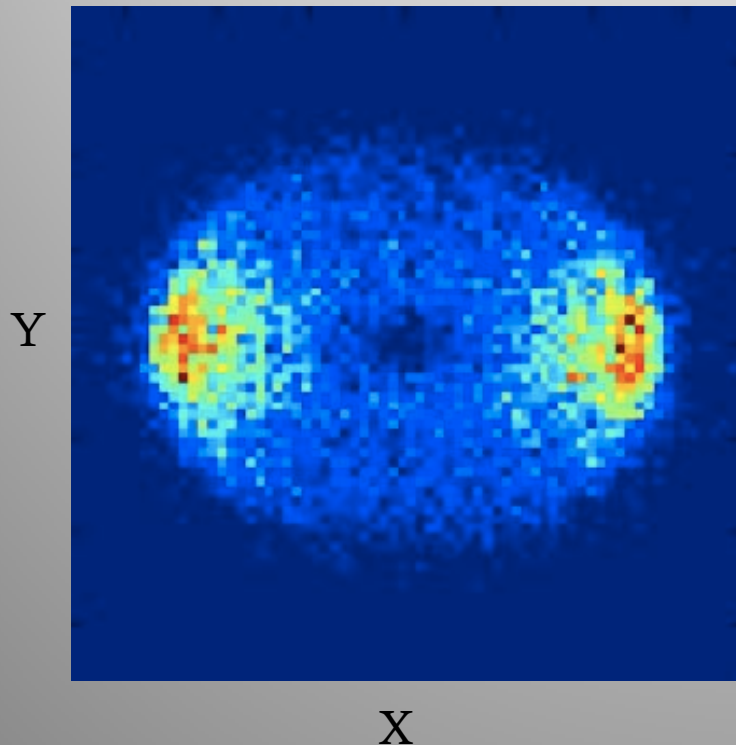


V_Z -- V_X Diagram of the X shape



Particles with Banana Orbit

- More evidence from the spatial distribution of particles with banana orbits in our simulation (Model 2)



Summary

- The buckling process thickens the bar to form an inner B/PS bulge, which has distinct kinematic properties compared to the outer thin component of the bar.
 - The **inner** component: **dynamically hot, small average velocity**
 - The **outer** component: **dynamically cold, large average velocity**
- Surface density profile along the bar major axis can be well described with a single Sérsic function, with stronger buckled bar a larger Sérsic index.
 - **No evidence for two components in the density profile**
- Relative contributions of the inner bifurcated and outer unperturbed $x1$ orbits produce the observed peanut shape, which also depends on the strength of the buckling.
- The particle motions within the X-shaped regions agree well with the banana-like orbits, which produces strong positive slope in V_Z -- V_X diagram.